M1.(a) an object with an escape velocity greater than the speed of light ✓ Ignore references to singularity and density etc. Allow gravity so strong light cannot escape.

(b) mass of black hole = $1 \times 10^{10} \times 1.99 \times 10^{30} = 2 \times 10^{40}$ kg \checkmark *M* correct for the first mark

Use of R = 2GM / c^2 = 2 × 6.67 × 10⁻¹¹× 2 × 10⁴⁰ / (3.00 × 10⁸)² = 3 × 10¹³ m \checkmark allow 2.9 or 2.95 etc. Final answer correct for the second mark. Allow ce for the mass. No sf penalty.

(c) V= Hd $v (in kms^{-1}) = 6300$ D (in MPc) = $3.3 \times 10^8 / 3.26 \times 10^6$ = 101 🗸 $H = v / d = 6300 / 101 = 62 \text{ kms}^{-1} \text{ Mpc}^{-1}$ Alternatively. Age of universe = 1/H= D/v $= 3.3 \times 10^8 \times 9.47 \times 10^{15} \checkmark / 6.3 \times 10^6 \checkmark$ $= 5.0 \times 10^{17} \text{ s}$ age of Universe = 1 / H = 1/62= 1.6 × 10⁻² Mpc s km⁻¹ $= 1.6 \times 10^{-2} \times 3.1 \times 10^{16} \times 10^{6} / 10^{3}$ = 5.0 × 10¹⁷ s ✓ The first mark is for calculating D, the second for substituting correctly to find H The third is for determining 1 / H in seconds. If other value of H used, 1 mark max.

1

3

M2. (a)	Gives the <u>ratio</u> of the (recessional) velocity (of galaxies) to distance from Earth Accept equation with terms defined not v depends on d, the relationship between them, shows the relationship between them				
		B1	1		
(b)	<i>d</i> changed to Mpc (2.45 × 10²) or 1.8 × 10⁴ / their attempt to convert distance <i>Or d change to m and v to m</i> s⁻¹				
		B1			
	(<i>H</i> =) 73.35 or 73.47 seen to at least 3 sf				
		B1	2		
(c)	(i) $T = 1 / H \text{ or } H = 2.4 \times 10^{-18} \text{ s seen}$ e.g. 3.08 × 10 ⁻¹⁹ / 73				
		C1			
	Value in s calculated (4.2×10^{17})				
		A1			
	Correct conversion to years 1.3 × 10 ¹⁰ Allow their value in s				
		B1	3		
	(ii) Universe is expanding at constant / steady <u>rate</u>				
		B1	1		

[7]

 M3.(a) (i) Similarity both would appear the same brightness As the apparent magnitudes are the same ✓
 Description and explanation needed for mark. Any references to same size gets zero for 1st mark.

> Difference Kocab would appear orange / red, Polaris yellow / white Due to their spectral classes / different temperatures ✓ Allow different colours + ref to spectral class for second mark If colour named, should be correct.

Polaris is further from Earth:
 Alternative:
 Polaris hotter and same size

Both stars same size and Polaris is hotter \checkmark

As P = σ AT⁴ Hence, Polaris has brighter absolute magnitude / is intrinsically brighter

Same A, would mean that Polaris has greater power output.

Polaris must be further from Earth to appear same brightness as Kocab. ✓ Same apparent brightness, therefore Polaris is further away.

(b) (i) v = Hd

- v = 0.025 × 3 × 10⁵ = 7.5 × 10³ km s⁻¹ ✓ 1⁵ mark is for calculating v
- d = 340 × 10^e l yr = 340 / 3.26 Mpc = 104 Mpc ✓ 2nd mark is for working out d in Mpc

 $H = 7.5 \times 10^{3} / 104 = 72 \text{ kms}^{-1} \text{ Mpc}^{-1} \checkmark$ 3rd mark is for calculating H in the correct unit.

3

(ii) Age of Universe = 1 / H 1st mark is for the equation

> = 0.014 × 10⁶ × 3.26 × 9.5 × 10¹⁵ / 1000 2nd is for the answer with working

= 4.3 × 10¹⁷ seconds

(= 13.6 billion years)

Unit consistent with calculation.

3rd is for a time unit consistent with their answer / working

3

1

2

M4.	(a)	(i)	increase in wavelength (of em radiation) due to relative		
	recessive velocity between observer and source 💉				

- (ii) use of v = Hd to give v = $65 \times 25 \sqrt{}$ = 1.6×10^3 (km s⁻¹) $\sqrt{}$
- (b) (i) all type 1a supernovae have same **peak** absolute magnitude ✓
 apparent magnitude can be measured ✓
 ref to m-M log (d/10) or inverse square law ✓
 - (ii) use of m-M = 5 log (d/10) gives $12.9 - (-19.3) = 5 \log (d/10) \sqrt{2}$ $\log (d/10) = 6.44$ $d = 27.5 (Mpc) \sqrt{2}$

(c) to make the accepted value for the distance more reliable 🗸

1

M5. (a) (use of
$$\frac{\Delta \hat{\lambda}}{\hat{\lambda}} = -\frac{\nu}{c}$$
 gives) $\frac{(660.86 - 656.28)}{656.28} = (-)\frac{\nu}{3.0 \times 10^8}$ (1)
 $\nu = (-)2094 \text{ km s}^{-1}$ (1)

(b) graph to show: correct plotting of points (1) straight line through origin (1) $H = \frac{v}{d} = \text{gradient} = 70 \text{ km s}^{-1} \text{ Mpc}^{-1} (1)$

20×10⁶

(must show evidence of use of graph in calculation)

3

[5]

2

(iii) (use of
$$\frac{\Delta \hat{\lambda}}{\hat{\lambda}} = -\frac{\nu}{c}$$
 gives) $\Delta \lambda = \frac{1000 \times 10^3}{3 \times 10^8} \times 656.3$ (nm) = 2.19 (nm) (1)
(allow C.E. for value of *v* from (ii)
 $\lambda_{galaxy} = 656.3 + 2.19 = 658.5$ nm (1)

(b) for the furthest point of the Universe,
$$d = \frac{c}{H}$$
 (1)

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age of Universe = $\frac{d}{c} = \frac{1}{H}$ (1) [or use of v = Hd and $t = \frac{d}{v}$ (1) if all started from same point $t = age of Universe = \frac{1}{H}$ (1)] assumption: that *H* remains constant

[7]

3